

Quantum Mechanics II

Quantum \rightarrow Classical Physics

**The “Classical Limit”;
Ehrenfest’s Theorem;
The Phase \rightarrow Hamilton-Jacobi Theorem**

Tristan Hübsch

Department of Physics and Astronomy, Howard University, Washington DC

<http://physics1.howard.edu/~thubsch/>

Quantum → Classical Physics

The “Classical Limit”

- Often dubbed the “ $\hbar \rightarrow 0$ ” limit
- Strictly, “ $1.055 \times 10^{-34} \text{ J}\cdot\text{s} \rightarrow 0$ ” is as *meaningless* as “ $1 \rightarrow 0$ ”
- A typical classical event:
 - A piece of chalk (0.01 kg) thrown against the blackboard (2 m away)
 - flies $\sim 1/10 \text{ s}$ so, $v \sim (2 \div 1/10 = 20) \text{ m/s}$ so, $KE \sim 1/2(0.01)(20)^2 = 2 \text{ J}$
 - so, Hamilton’s action $= \int dt (KE - PE) \sim (1/10)(2) = 0.2 \text{ J}\cdot\text{s}$
 vs. $\hbar = 1.055 \times 10^{-34} \text{ J}\cdot\text{s}$
 - A pirouetting ballerina (60 kg), spinning at 1 Hz, *i.e.*, $\omega = 2\pi \cdot 1 \text{ s}^{-1}$
 - $KE_{\text{rot}} \sim 1/2(1/2(60)(0.20)^2)(2\pi)^2 = 23.69 \text{ J}$, Hamilton’s action $\sim (1)(23.69) \text{ J}\cdot\text{s}$
 (axially rotating cylinder) *vs.* $\hbar = 1.055 \times 10^{-34} \text{ J}\cdot\text{s}$
- Instead, test whether $(\hbar/S_i) \ll 1$, for all S_i **commensurate** with \hbar
- If true, classical; if false (for ↑ even one S_i), quantum. Note: This has little to do with size, or “amounts” of position, momentum, energy, time... *per se* !!!

key ratios!!!

Quantum → Classical Physics

Ehrenfest's Theorem

- The claim: Expectation values of quantum operators obey classical dynamical equations
- Expectation values are state-dependent
 - represent average trajectories, energies, momenta, ...
 - need not represent any particular trajectory, energy, momentum, ...
 - just like there exists no particular USA family with 2.3 children
- A (quantum) state thus \Leftrightarrow an ensemble of classical states
- Classical equation of motion:

$$\frac{dR}{dt} = \frac{\partial R}{\partial t} + \{H, R\}_{\text{PB}} = \frac{\partial R}{\partial t} + \left(\frac{\partial H}{\partial q^i} \frac{\partial R}{\partial p_i} - \frac{\partial R}{\partial q^i} \frac{\partial H}{\partial p_i} \right)$$

$$\frac{d\langle \hat{R} \rangle}{dt} = \frac{\partial \langle \hat{R} \rangle}{\partial t} + \{ \langle \hat{H} \rangle, \langle \hat{R} \rangle \}_{\text{PB}} = \frac{\partial \langle \hat{R} \rangle}{\partial t} + \left(\frac{\partial \langle \hat{H} \rangle}{\partial q^i} \frac{\partial \langle \hat{R} \rangle}{\partial p_i} - \frac{\partial \langle \hat{R} \rangle}{\partial q^i} \frac{\partial \langle \hat{H} \rangle}{\partial p_i} \right)$$

Quantum → Classical Physics

Ehrenfest's Theorem

Simple example: $\hat{H} = \frac{\hat{P}^2}{2M} + W(\hat{Q}) \quad \frac{\partial \hat{P}}{\partial t} = 0 = \frac{\partial \hat{Q}}{\partial t}$

Calculate:

$$\frac{d\hat{Q}}{dt} = \frac{i}{\hbar} [\hat{H}, \hat{Q}] = \frac{\hat{P}}{M}$$

$$\frac{d\hat{P}}{dt} = \frac{i}{\hbar} [\hat{H}, \hat{P}] = -W'(\hat{Q})$$

These are the (quantum, Heisenberg) equations of motion

Compute the expectation values of these:

$$\frac{d\langle \hat{Q} \rangle}{dt} = \frac{\langle \hat{P} \rangle}{M} \quad \frac{d\langle \hat{Q} \rangle}{dt} = -\langle W'(\hat{Q}) \rangle \stackrel{?}{=} -W'(\langle \hat{Q} \rangle)$$

must be verified

as claimed

In systems w/o EM interaction, the proof reduces to proving

$$\langle W'(\hat{Q}) \rangle \stackrel{?}{=} W'(\langle \hat{Q} \rangle)$$

Quantum → Classical Physics

Ehrenfest's Theorem

● In general,

$$\frac{d\hat{R}}{dt} = \frac{\partial \hat{R}}{\partial t} + \frac{i}{\hbar} [\hat{H}, \hat{R}]$$

$$\frac{dR}{dt} = \frac{\partial R}{\partial t} + \{H, R\}_{\text{PB}}$$

How?

● Expand:

$$\hat{H} = \langle \hat{H} \rangle + \hat{\mathcal{H}} \quad \hat{R} = \langle \hat{R} \rangle + \hat{\mathcal{R}} \quad \frac{i}{\hbar} [\hat{H}, \hat{R}] = F(\hat{Q}, \hat{P}, \hat{R}, \dots)$$

$$F(\hat{Q}, \hat{P}, \hat{R}, \dots) = F(\langle \hat{Q} \rangle + \hat{\mathcal{Q}}, \langle \hat{P} \rangle + \hat{\mathcal{P}}, \langle \hat{R} \rangle + \hat{\mathcal{R}}, \dots)$$

$\hat{Q}-\hat{P}-\hat{R}-\dots$ coupling! ...a complex issue!

$$= F(\langle \hat{Q} \rangle, \langle \hat{P} \rangle, \langle \hat{R} \rangle, \dots) + \left[\frac{\partial F}{\partial \langle \hat{R} \rangle} \right]_0 \hat{\mathcal{R}} + \left[\frac{1}{2} \frac{\partial^2 F}{\partial \langle \hat{R} \rangle^2} \right]_0 \hat{\mathcal{R}}^2 + \dots$$

“quite often”

+ terms coupling with $\hat{\mathcal{Q}}, \hat{\mathcal{P}}, \dots$

$$\stackrel{?}{=} \{ \langle \hat{H} \rangle, \langle \hat{R} \rangle \}_{\text{PB}} + \dots$$

● It is possible (albeit tedious) to pursue this

● “Background” expansion: vevs = classical background

Quantum → Classical Physics

Ehrenfest's Theorem

- In general,
- Expand:

$$\frac{dR}{dt} = \frac{\partial R}{\partial t} + \{H, R\}_{\text{PB}}$$

$$\frac{d\hat{R}}{dt} = \frac{\partial \hat{R}}{\partial t} + \frac{i}{\hbar} [\hat{H}, \hat{R}] \quad \text{How?}$$

$$\hat{H} = \langle \hat{H} \rangle + \hat{\mathcal{H}} \quad \hat{R} = \langle \hat{R} \rangle + \hat{\mathcal{R}} \quad \frac{i}{\hbar} [\hat{H}, \hat{R}] = F(\hat{Q}, \hat{P}, \hat{R}, \dots)$$

$$F(\hat{Q}, \hat{P}, \hat{R}, \dots) = F(\langle \hat{Q} \rangle + \hat{\mathcal{Q}}, \langle \hat{P} \rangle + \hat{\mathcal{P}}, \langle \hat{R} \rangle + \hat{\mathcal{R}}, \dots)$$

$\hat{Q}-\hat{P}-\hat{R}-\dots$ coupling! *...a complex issue!*

$$= F(\langle \hat{Q} \rangle, \langle \hat{P} \rangle, \langle \hat{R} \rangle, \dots) + \left[\frac{\partial F}{\partial \langle \hat{R} \rangle} \right]_0 \hat{\mathcal{R}} + \left[\frac{1}{2} \frac{\partial^2 F}{\partial \langle \hat{R} \rangle^2} \right]_0 \hat{\mathcal{R}}^2 + \dots$$

"quite often"

+ terms coupling with $\hat{\mathcal{Q}}, \hat{\mathcal{P}}, \dots$

$$\stackrel{?}{=} \{ \langle \hat{H} \rangle, \langle \hat{R} \rangle \}_{\text{PB}} + \dots$$

- It is possible (albeit tedious) to pursue this
 - "Background" expansion: vevs = classical background
 - about which quantum observables fluctuate

Quantum → Classical Physics

Ehrenfest's Theorem

● In general,

$$\frac{d\hat{R}}{dt} = \frac{\partial \hat{R}}{\partial t} + \frac{i}{\hbar} [\hat{H}, \hat{R}]$$

$$\frac{dR}{dt} = \frac{\partial R}{\partial t} + \{H, R\}_{\text{PB}}$$

How?

● Expand:

$$\hat{H} = \langle \hat{H} \rangle + \hat{\mathcal{H}} \quad \hat{R} = \langle \hat{R} \rangle + \hat{\mathcal{R}} \quad \frac{i}{\hbar} [\hat{H}, \hat{R}] = F(\hat{Q}, \hat{P}, \hat{R}, \dots)$$

$$F(\hat{Q}, \hat{P}, \hat{R}, \dots) = F(\langle \hat{Q} \rangle + \hat{\mathcal{Q}}, \langle \hat{P} \rangle + \hat{\mathcal{P}}, \langle \hat{R} \rangle + \hat{\mathcal{R}}, \dots)$$

$\hat{Q}-\hat{P}-\hat{R}-\dots$ coupling! ...a complex issue!

$$= F(\langle \hat{Q} \rangle, \langle \hat{P} \rangle, \langle \hat{R} \rangle, \dots) + \left[\frac{\partial F}{\partial \langle \hat{R} \rangle} \right]_0 \hat{\mathcal{R}} + \left[\frac{1}{2} \frac{\partial^2 F}{\partial \langle \hat{R} \rangle^2} \right]_0 \hat{\mathcal{R}}^2 + \dots$$

“quite often”

+ terms coupling with $\hat{\mathcal{Q}}, \hat{\mathcal{P}}, \dots$

$$\stackrel{?}{=} \{ \langle \hat{H} \rangle, \langle \hat{R} \rangle \}_{\text{PB}} + \dots$$

● It is possible (albeit tedious) to pursue this

● “Background” expansion: vevs = classical background

● about which quantum observables fluctuate

● ...aiming to reproduce the classical equations to 0th order + a nonlinearly coupled system of quantum fluctuations...

Quantum → Classical Physics

Anomalies

$$\frac{dR}{dt} = \frac{\partial R}{\partial t} + \{H, R\}_{\text{PB}}$$

- Actually, the issue

$$\frac{d\hat{R}}{dt} = \frac{\partial \hat{R}}{\partial t} + \frac{i}{\hbar} [\hat{H}, \hat{R}]$$
 commutator \leftrightarrow Poisson bracket
- the correspondence
- ...turns out to be very important!
- In classical physics, observables will satisfy some Poisson-bracket algebra: $\{A_i, A_j\}_{\text{PB}} = f_{ij}^k A_k$ (as do rotations)
- The corresponding quantum operators need not do so.
- Denote the (“quantization”) assignment as $A_i \mapsto \hat{A}_i := \pi(A_i)$
- Then compute: $\Delta_{ij} := \pi(\{A_i, A_j\}_{\text{PB}}) - \frac{i}{\hbar} [\pi(A_i), \pi(A_j)]$
- If nonzero, then the quantization assignment $\pi(\dots)$ does not preserve the classical relationship between observables A_i .
- This could range from irrelevant (“ignorable” A_i ’s 🙈🙈🙈)
- to disastrous: gauge symmetry (EM charge conservation 🤯🆘)

Quantum → Classical Physics

Ehrenfest's Theorem

● In the “typical” cases, where $\langle F(\hat{Q}) \rangle = F(\langle \hat{Q} \rangle) + \dots$ negligible

● ...so Ehrenfest's theorem holds for position / momentum equations,

● ...Ehrenfest's theorem is not guaranteed to hold for all observables

● Also, Ehrenfest's theorem is *insufficient* guarantee:

● *Some* systems satisfy Ehrenfest's theorem, even though they are manifestly non-classical

$$\langle W'(\hat{Q}) \rangle \not\equiv W'(\langle \hat{Q} \rangle) \quad \text{although its energy levels are } \textit{quantized}$$

for LHO

● Also, Ehrenfest's theorem is *not necessary* a guarantee:

● *Some* systems behave classically although they do not satisfy Ehrenfest's theorem

● A particle in a box (impenetrable walls) oscillates with $k_n = n\pi/L$, energy $E_n = n^2\hbar^2\pi^2/2mL^2$, frequency $\omega_n = E_n/\hbar = n^2\hbar\pi^2/2mL^2$ and period $T_n = 2\pi/\omega_n = (4mL^2/\hbar\pi)/n^2$ — which is *nothing* like $T_{cl} = 2L/v$?!

Yet, the average distribution of both cl. and q. particles are constant & flat...

Quantum → Classical Physics

Bohr's Complementarity Principle

- Note: orbital angular momentum $\sim \ell \cdot \hbar \lll$ macroscopic value
- Bohr: in the limit $n \rightarrow \infty$, quantum should resemble classical

WKB:

$$\psi_{\text{WKB}}^{(\pm)}(x) = \frac{C_{\pm}}{\sqrt{k(x)}} e^{\pm i \int d\zeta k(\zeta)} \quad \overline{|\psi_{\text{WKB}}^{(\text{st.w.})}(x)|^2} = \frac{|C_{\pm}|^2}{k(x)} \frac{1}{2}$$

- The classical probability density is proportional to the time the particle spends within Δx , i.e., $\Delta t / \Delta x = 1/v = 1/(\hbar k/M) \propto 1/k$

H-atom:

$$\langle r^2 \rangle - \langle r \rangle^2 \xrightarrow{\ell \rightarrow n-1} \left(\frac{a_0}{2}\right)^2 n^2 (1 + 2n)$$

$$\frac{\langle r^2 \rangle - \langle r \rangle^2}{\langle r^2 \rangle} \xrightarrow{\ell \rightarrow n-1} \frac{1}{2} \frac{1}{1+n} \xrightarrow{n \rightarrow \infty} 0$$

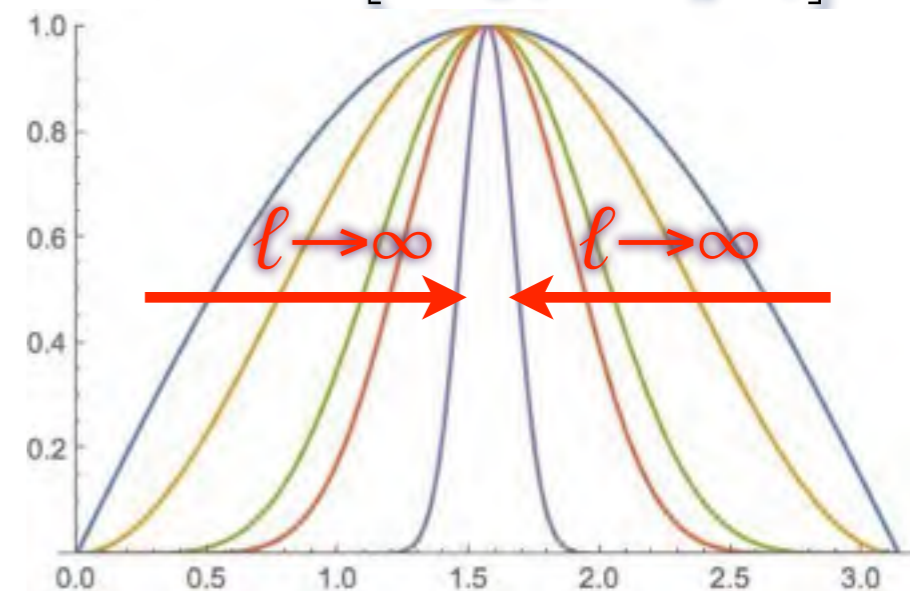
- Angular dependence: $|Y_{\ell}^{\ell}(\theta, \phi)|^2 \propto \sin^{2\ell}(\theta)$

- peaks @ $\theta = \pi/2$

equatorial orbits

$$\langle r^2 \rangle = n^4 a_0^2 \left[1 + \frac{3}{2} \left(1 - \frac{\ell(\ell+1) - \frac{1}{3}}{n^2} \right) \right],$$

$$\langle r \rangle = n^2 a_0 \left[1 + \frac{1}{2} \left(1 - \frac{\ell(\ell+1)}{n^2} \right) \right],$$



Quantum → Classical Physics

The Phase → Hamilton-Jacobi Theorem

- Inspired by the WKB approximation, consider

$$\Psi(\vec{r}, t) = A(\vec{r}, t) e^{iS(\vec{r}, t)/\hbar}$$

- for which the Schrödinger equation produces

$$-\frac{\partial S}{\partial t} = \frac{(\vec{\nabla} S)^2}{2M} + W - \frac{\hbar^2 \vec{\nabla}^2 A}{2M A}$$

$$\frac{\partial A}{\partial t} = -\frac{1}{2M} [\vec{\nabla} \cdot (A \vec{\nabla} S) + (\vec{\nabla} A) \cdot (\vec{\nabla} S)]$$

$$W_Q := -\frac{\hbar^2 \vec{\nabla}^2 A}{2M A}$$

Does NOT always vanish for $\hbar \rightarrow 0$

- Defining $P := A^2$, this becomes:

$$0 = \frac{\partial S}{\partial t} + \frac{(\vec{\nabla} S)^2}{2M} + (W + W_Q)$$

Hamilton-Jacobi equation, with $W \rightarrow W + W_Q$

$$0 = \frac{\partial P}{\partial t} + \frac{1}{M} \vec{\nabla} \cdot (P \vec{\nabla} S) \quad \vec{J} := (P \vec{\nabla} S) / M$$

probability density continuity equation

Quantum Mechanics II

*Now, go forth and
calculate!!!*

Tristan Hübsch

Department of Physics and Astronomy, Howard University, Washington DC

<http://physics1.howard.edu/~thubsch/>