

Quantum Mechanics II

Time-Dependent Physics

**Energy-Time Indeterminacy;
Quantum Beats;
Kaons and Neutrinos**

Tristan Hübsch

Department of Physics and Astronomy, Howard University, Washington DC

<http://physics1.howard.edu/~thubsch/>

Time-Dependent Phenomena

Energy-Time Indeterminacy

- Heisenberg's indeterminacy relations $\Delta_Q \Delta_P \geq \frac{1}{2}\hbar$
- Often, as “dual variables in Fourier transformation”
 - Cartesian coordinates, $\mathcal{F}: f(x) \rightarrow F(k_x)$, and $k_x := p_x/\hbar$. *...however...*
 - Typically, $\mathcal{F}: f(t) \rightarrow F(\omega)$, and $\omega := (E_2 - E_1)/\hbar$, and so $\omega \neq E/\hbar$
 - Ballentine: energy may be shifted $E \rightarrow E + E_0$, with $E_0 = \text{const.}$ arbitrary
 - But, so can $p \rightarrow p + p_0$: for a fixed mass, this is just a Galilean boost.
- Sometimes, as “canonically conjugate variables”
 - No general precise derivation (dimensional analysis does check out)
 - Not general enough...

Time-Dependent Phenomena

General Indeterminacy Relations

● Robertson ('29), Schrödinger ('30), Jackiw & Carruthers+Nieto ('68):

● Given two Hermitian operators, define the third one as

$$\hat{C} := -i[\hat{A}, \hat{B}] = -i[(\hat{A} - \langle \hat{A} \rangle), (\hat{B} - \langle \hat{B} \rangle)]$$

$$\begin{aligned} 0 &\leq \left\langle \left([\hat{A} - \langle \hat{A} \rangle] - i\zeta[\hat{B} - \langle \hat{B} \rangle] \right)^\dagger \left([\hat{A} - \langle \hat{A} \rangle] - i\zeta[\hat{B} - \langle \hat{B} \rangle] \right) \right\rangle \\ &= \left\langle |\hat{A} - \langle \hat{A} \rangle|^2 \right\rangle - i\zeta \left\langle [\hat{A} - \langle \hat{A} \rangle, \hat{B} - \langle \hat{B} \rangle] \right\rangle + \zeta^2 \left\langle |\hat{B} - \langle \hat{B} \rangle|^2 \right\rangle \\ &= (\Delta_A)^2 + \zeta \langle \hat{C} \rangle + \zeta^2 (\Delta_B)^2 \end{aligned}$$

● True for all ζ , this is true for $\min(\zeta) = -\langle \hat{C} \rangle / 2(\Delta_B)^2$

$$= (\Delta_A)^2 - \frac{\langle \hat{C} \rangle^2}{2(\Delta_B)^2} + \frac{\langle \hat{C} \rangle^2}{4(\Delta_B)^2} = (\Delta_A)^2 - \frac{\langle \hat{C} \rangle^2}{4(\Delta_B)^2}$$

$$(\Delta_A)^2 \geq \frac{\langle \hat{C} \rangle^2}{4(\Delta_B)^2} \quad (\Delta_A)^2 (\Delta_B)^2 \geq \frac{1}{4} \langle \hat{C} \rangle^2$$

≥ 0
 $(\hat{C}^\dagger = \hat{C})$

$$\Delta_A \Delta_B \geq \frac{1}{2} \left| \langle [\hat{A}, \hat{B}] \rangle \right|$$

state-dependent

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 - No general precise derivation (dimensional analysis does check out)
 - Not general enough...
- Nonrelativistic Quantum Mechanics
 - coordinates are eigenvalues (expectation values) of Hermitian operators
 - time is not.

Time is a parameter
...on which everything else depends

Time-Dependent Phenomena

Energy-Time Indeterminacy

- Wolfgang Pauli's "TH-theorem" (1933)
- Suppose time was the eigenvalue of a Hermitian operator, canonically conjugate to the Hamiltonian

$$[\hat{T}, \hat{H}] = i\hbar \quad \hat{U}_\varepsilon := \exp\{-i\varepsilon\hat{T}/\hbar\}$$

$$[\hat{H}, \hat{U}_\varepsilon] = [\hat{H}, \hat{T}] \frac{\partial \hat{U}_\varepsilon}{\partial \hat{T}} = (-i\hbar)\hat{U}_\varepsilon(-i\varepsilon/\hbar) = -\varepsilon\hat{U}_\varepsilon$$

- Then,

$$\hat{H}|E\rangle = E|E\rangle \quad \hat{H}(\hat{U}_\varepsilon|E\rangle) = (\hat{U}_\varepsilon\hat{H} - \varepsilon\hat{U}_\varepsilon)|E\rangle = (E - \varepsilon)(\hat{U}_\varepsilon|E\rangle)$$

Time-Dependent Phenomena

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$$\hat{U}_\varepsilon|E\rangle = |E - \varepsilon\rangle \quad \varepsilon \rightarrow \infty \Rightarrow \exists \text{ normalized state with } E \rightarrow -\infty$$

\Rightarrow no ground-state with definite energy!

- Thus, time cannot be the eigenvalue of a Hermitian operator.

Time-Dependent Phenomena

Energy-Time Indeterminacy

- So, what *should* a relationship such as “ $\Delta_E \Delta_\tau \geq i\hbar$ ” mean?
 - It is not time itself that is observed, but sequential variations in some other observable...
 - ...which then serves as a clock.

- Consider then an observable that is not stationary

$$[\hat{H}, \hat{R}] \neq 0 \quad \Delta_R \Delta_H \geq \frac{1}{2} |\langle [\hat{H}, \hat{R}] \rangle| = \frac{\hbar}{2} \left| \frac{d\langle \hat{R} \rangle}{dt} \right|$$

S. Mandelstam & I. Tamm

- Then, define:

$$\tau_R := \Delta_R \left| \frac{d\langle \hat{R} \rangle}{dt} \right|^{-1} \quad \tau_R \Delta_H \geq \frac{1}{2} \hbar$$

- This τ_R serves as a characteristic time (period) of any phenomenon in which variations in the observable R serve as a clock. Whence “clock-observable.”

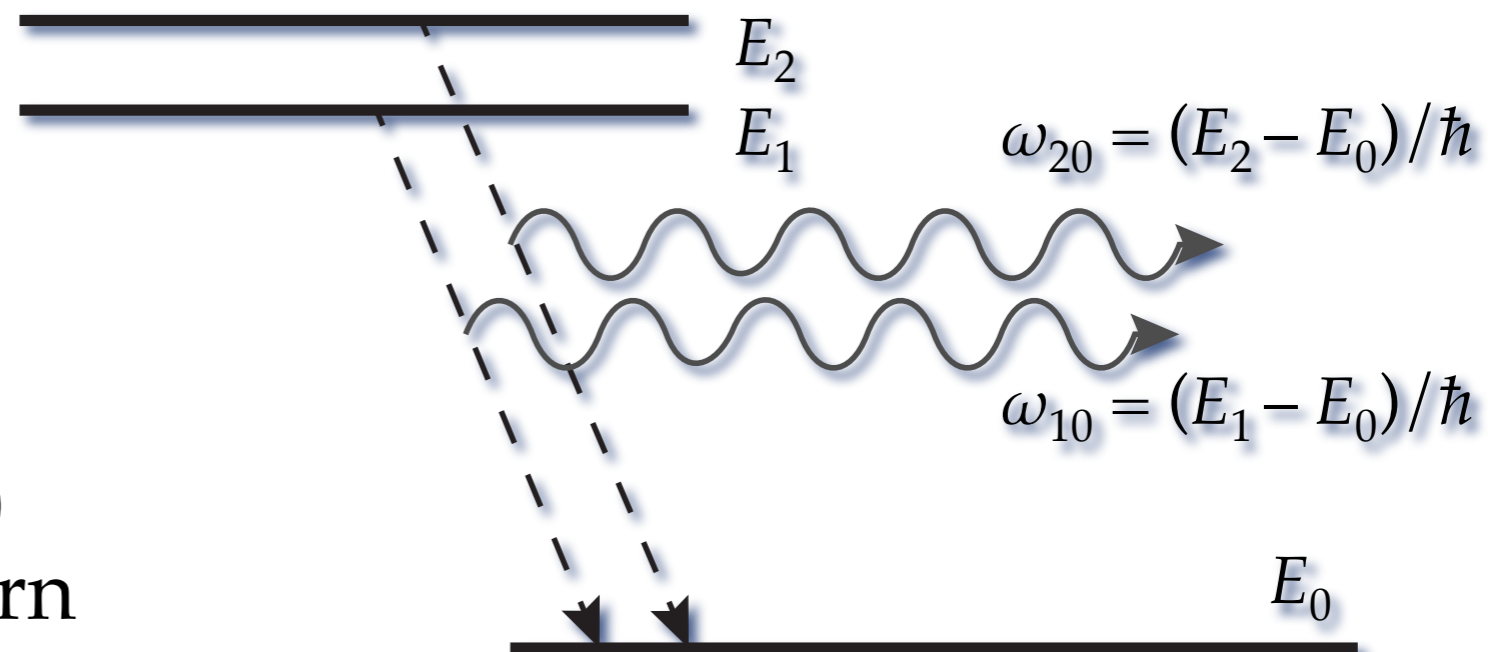
Time-Dependent Phenomena

Quantum Beats

- Recall the spin- $1/2$ system (particle) with $\mu = \gamma S$
 - ...gaining/losing $1/2\gamma\hbar\mathbf{B}\cdot\boldsymbol{\sigma}$ energy in the \mathbf{B} -field
 - ...with μ precessing with frequency $\omega = \gamma |\mathbf{B}|$
- A linear combination of the two states is not stationary

$$\hat{U}_t \begin{bmatrix} c_+ \\ c_- \end{bmatrix} = e^{-i\omega t \sigma_z} \begin{bmatrix} c_+ \\ c_- \end{bmatrix} = \begin{bmatrix} c_+ e^{-i\omega t} \\ c_- e^{+i\omega t} \end{bmatrix}$$

- Similar result for all 2-state systems
- Suppose this linear combination decays into a lower state
- The two emitted photons interfere
- Int. \propto Prob. $\propto \cos(\omega_{21} t)$ modulated decay pattern



Time-Dependent Phenomena

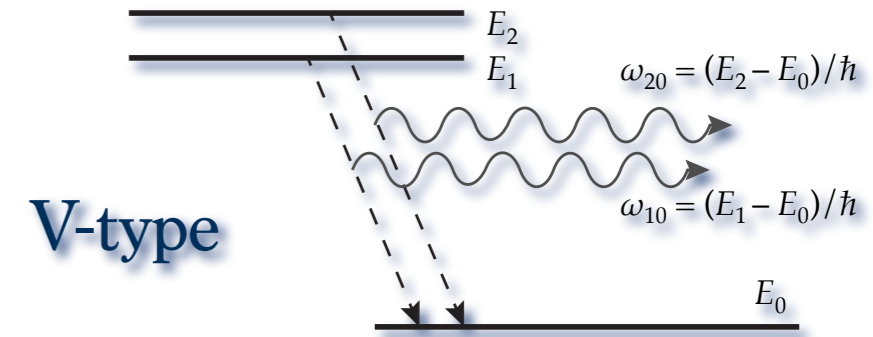
Quantum Beats

- Treating the EM radiation classically,

$$I \propto \sin(\omega_{20}t) + \sin(\omega_{10}t)$$

$$= 2 \sin\left(\frac{\omega_{20} + \omega_{10}}{2}t\right) \sin\left(\frac{\omega_{20} - \omega_{10}}{2}t\right)$$

“carrier”
“AM” = beats

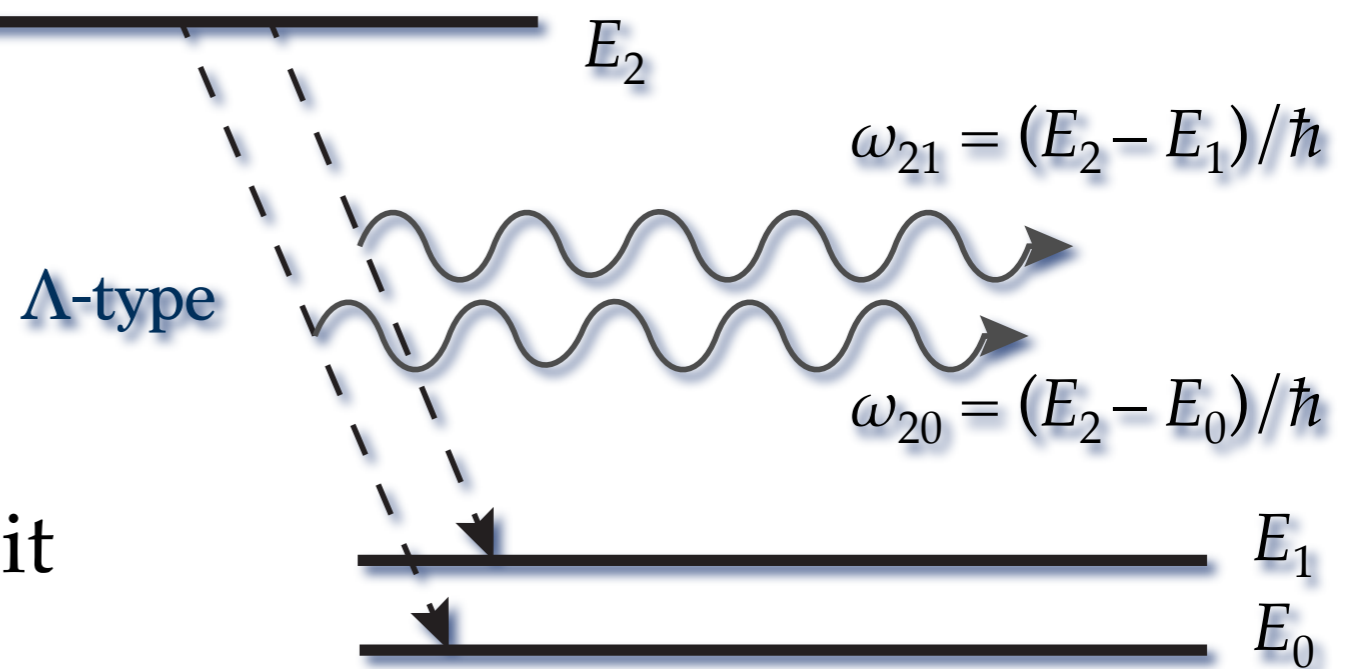


- A similar effect should also exist in the flipped situation...

- ...while the EM radiation is treated classically

- But, V-type atoms do exhibit beats as predicted...

- ...while Λ -type atoms do not.



Time-Dependent Phenomena

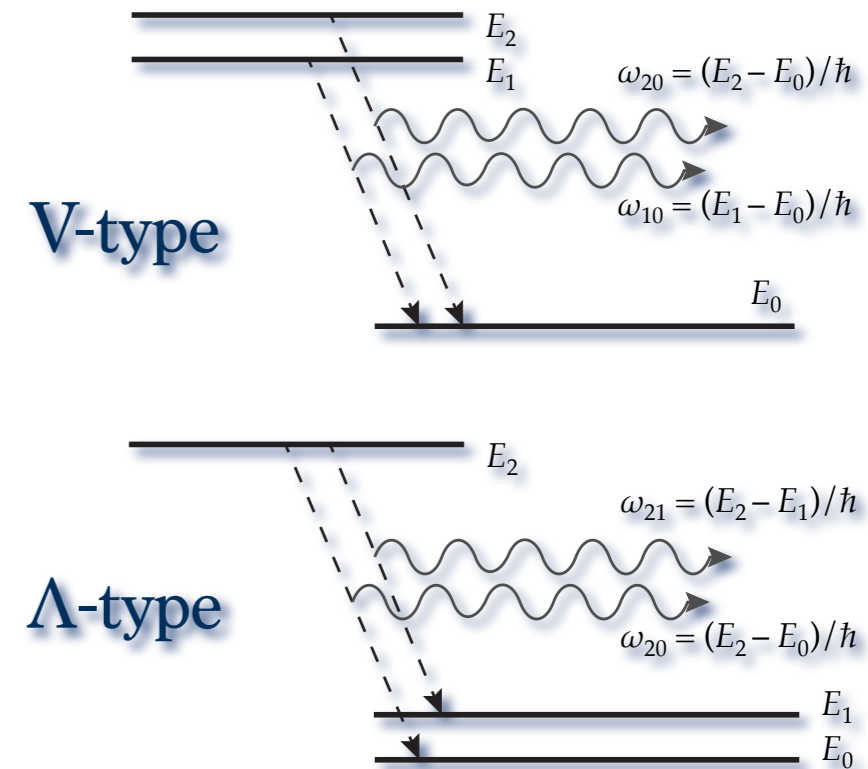
Quantum Beats

- The full quantum description of the two cases uses the final states:

$$\alpha_1 |E_0; \omega_{10}\rangle + \alpha_2 |E_0; \omega_{20}\rangle$$

$$\beta_1 |E_1; \omega_{21}\rangle + \beta_2 |E_0; \omega_{20}\rangle$$

EM radiation



- Beats are caused by interference
- ...for which the probability amplitudes are

$$\langle E_0; \omega_{10} | \hat{H}_\gamma | E_0; \omega_{20} \rangle_{\text{V-type}}$$

$$\langle E_1; \omega_{21} | \hat{H}_\gamma | E_0; \omega_{20} \rangle_{\Lambda\text{-type}}$$

$$\underbrace{\langle E_0 | E_0 \rangle}_{=1} \cdot \langle \omega_{10} | \hat{H}_\gamma | \omega_{20} \rangle$$

beats

$$\underbrace{\langle E_1 | E_0 \rangle}_{=0} \cdot \langle \omega_{21} | \hat{H}_\gamma | \omega_{20} \rangle$$

~~beats~~

- Useful when E_1 and E_2 cannot be resolved experimentally

Time-Dependent Phenomena

Kaons

- Two neutral, spin-0 mesons
 - one decays into two pions, after 8.958×10^{-11} s = K_S $CP = +1$
 - the other into three pions, after 5.114×10^{-8} s = K_L $CP = -1$
 - although they are created in same collision processes

So, K_S and K_L are decay eigenstates

The creation eigenstates are $K_0 = (K_S + K_L)$ and $\bar{K}_0 = (K_S - K_L)$

Created 50%-50%, the ratio soon depletes

$$\frac{N(K_S)}{N(K_L)} = \frac{e^{-t/\tau_S}}{e^{-t/\tau_L}} = \exp \left\{ -\frac{t}{\tau_S} + \frac{t}{\tau_L} \right\} \approx \exp \left\{ -569.9 \frac{t}{\tau_L} \right\}$$

which drops to 1.447×10^{-5} after just 1 ns!

However, virtual decay-undecay processes can oscillate $K_S \leftrightarrow K_L$

...and yield two-pion decays even after many seconds of flight.

This regeneration violates CP conservation *needed for Big-Bang baryogenesis*

Time-Dependent Phenomena

Neutrino Oscillations

● β -decay: ${}_Z X \rightarrow {}_{Z\pm 1} Y + e^\pm$ (i.e., $n^0 \rightarrow p^+ + e^-$ or $p^+ \rightarrow n^0 + e^+$)
cannot satisfy both energy and momentum conservation

● W. Pauli (1930): a third (very light, neutral) particle
E. Fermi: "small neutron" = neutrino

● $n^0 \rightarrow p^+ + e^- + \bar{\nu}$ or $p^+ \rightarrow n^0 + e^+ + \nu$

● In the next 1-2 decades, cosmic ray sources:

$$\begin{aligned} \pi^- &\rightarrow \mu^- + \bar{\nu}_\mu \\ &\rightarrow (e^- + \bar{\nu}_e + \nu_\mu) + \bar{\nu}_\mu \end{aligned}$$

$$\begin{aligned} \pi^+ &\rightarrow \mu^+ + \nu_\mu \\ &\rightarrow (e^+ + \nu_e + \bar{\nu}_\mu) + \nu_\mu \end{aligned}$$

● There should be twice as many cosmic ν_μ 's than ν_e 's.

● Vertically, yes; $\approx 2:1$. Horizontally (along the horizon), no; $\approx 1:1$

Time-Dependent Phenomena

Neutrino Oscillations

- By 1938, Hans Bethe: “Carbon cycle” & “pp-process”
H → He fusion in stars such as the Sun
- ...with a detailed spectrum of neutrinos predicted
- ...of which only $\sim 1/3$ ν_e 's arrive to Earth (ν_μ 's not detected)
- Creation & detection = \hat{H}_I -eigenstates, ν_e and ν_μ
- Propagation/evolution = \hat{H}_0 -eigenstates, say “1” and “2”
 - ...these are also the mass-eigenstates
- Created as, say, ν_e :

$$e^{-it\hat{H}_0/\hbar} \left(|\nu_e; 0\rangle := |“1+2”; 0\rangle \right) = |“1+2”; t\rangle$$

$$= C_1 e^{-iE_1 t/\hbar} |1\rangle + C_2 e^{-iE_2 t/\hbar} |2\rangle$$

$$P_\alpha := \left| [\cos(\alpha)\langle 1| + \sin(\alpha)\langle 2|] |“1+2”; t\rangle \right|^2$$

$$= |C_1|^2 \cos^2(\alpha) + |C_2|^2 \sin^2(\alpha) + \sin(2\alpha) \Re [C_1 C_2^* e^{-i(E_1 - E_2)t/\hbar}]$$

Time-Dependent Phenomena

Neutrino Oscillations $P_\alpha := \left| [\cos(\alpha)\langle 1| + \sin(\alpha)\langle 2|] | \text{“1+2”}; t \rangle \right|^2$
 $= |C_1|^2 \cos^2(\alpha) + |C_2|^2 \sin^2(\alpha) + \sin(2\alpha) \Re [C_1 C_2^* e^{-i(E_1 - E_2)t/\hbar}]$

- So, if the neutrino was initially in the “opposite” linear combination, $-\sin(\alpha)|1\rangle + \cos(\alpha)|2\rangle$, (the $\alpha + \pi/2$ state)

$$P_{|\alpha + \frac{\pi}{2}\rangle \rightarrow |\alpha\rangle} = \sin^2(2\alpha) \sin^2\left(\frac{1}{2}\omega_{12}t\right), \quad \omega_{12} := \frac{E_1 - E_2}{\hbar}$$

- The neutrino oscillates

$$\begin{aligned} |\nu_e\rangle &:= |\alpha\rangle = \cos(\alpha)|1\rangle + \sin(\alpha)|2\rangle \\ |\nu_\mu\rangle &:= |\alpha + \frac{\pi}{2}\rangle = -\sin(\alpha)|1\rangle + \cos(\alpha)|2\rangle \end{aligned}$$

- ...provided:

- The two stationary states are not degenerate, $E_1 \neq E_2$, $\omega_{12} \neq 0$

- The interaction eigenstates are not equal to the stationary states, $\alpha \neq 0$

and $E_1 - E_2 = \sqrt{|\vec{p}|^2 c^2 + m_1^2 c^4} - \sqrt{|\vec{p}|^2 c^2 + m_2^2 c^4} \approx \frac{(m_1^2 - m_2^2)c^4}{2\bar{E}}$

and, there are really three distinct neutrinos...

Quantum Mechanics II

*Now, go forth and
calculate!!!*

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